Chapter 7

Brownian Motion: Fokker-Planck Equation

The Fokker-Planck equation is the equation governing the time evolution of the probability density of the Brownian particla. It is a second order differential equation and is exact for the case when the noise acting on the Brownian particle is Gaussian white noise. A general Fokker-Planck equation can be derived from the Chapman-Kolmogorov equation, but we also like to find the Fokker-Planck equation corresponding to the time dependence given by a Langevin equation.

The derivation of the Fokker-Planck equation is a two step process. We first derive the equation of motion for the probability density $4/\operatorname{varrho}(x, v, t)4$ to find the Brownian particle in the interval (x, x+dx) and (v, v+dv) at time t for one realization of the random force $\xi(t)$. We then obtain an equation for

$$P(x, v, t) = \langle \varrho(x, v, t) \rangle_{\xi}$$

i.e. the average of $\rho(x, v, t)$ over many realizations of the random force. The probability density P(x, v, t) is the macroscopically observed probability density for the Brownian particel.

7.1 Probability flow in phase-space

Let us obtain the probability to find the Brownian particle in the interval (x, x + dx) and (v, v + dv) at time t. We will consider the space of coordiantes $\mathbf{x} = (x, v)$. The Brownian particle is located in the infinitesimal ara dxdv with probability $\rho(x, v, t)dxdv$. The velocity of the particle at point (x, v) is given by $\dot{\mathbf{x}} = (\dot{x}, \dot{v})$ and the current density is $\dot{\mathbf{x}}\rho$. Since the Brownian particle must lie somewhere in the phase-space $-\infty < x < \infty, \infty < v < \infty$ we have the condition

$$\int_{-\infty}^{\infty} \mathrm{d}x \int_{-\infty}^{\infty} \mathrm{d}v \varrho(x, v, t) = 1$$

Let us now consider a finite area, or volume, V_0 in this space. Since the Brownian particle cannot be destroyed a change in the probability contained in V_0 must be due to a flow of probability through the surface S_0 surrounding V_0 . Thus

$$\frac{\mathrm{d}}{\mathrm{d}t} \int \int_{V_0} \mathrm{d}x \mathrm{d}v \varrho(x, v, t) = -\int_{S_0} \varrho(x, v, t) \dot{\boldsymbol{x}} \cdot \mathrm{d}\boldsymbol{S}$$

We can now use Gauss theorem to change the surface integral into a volume integral.

$$\int \int_{V_0} \mathrm{d}x \mathrm{d}v \frac{\partial}{\partial t} \varrho(x, v, t) = -\int \int_{V_0} \mathrm{d}x \mathrm{d}v \nabla \cdot (\dot{x} \varrho(x, v, t))$$

Since V_0 is fixed and arbitrary we find the continuity equation

$$\frac{\partial}{\partial t}\varrho(x,v,t) = -\nabla \cdot (\dot{x}\varrho(x,v,t)) = -\frac{\partial}{\partial x} \left(\dot{x}\varrho(x,v,t) \right) - \frac{\partial}{\partial v} \left(\dot{v}\varrho(x,v,t) \right)$$
(7.1)

This is the continuity equation in phase-space which just state that probability is conserved.

7.2 Probability flow for Brownian particle

In order to write (7.1) explicitly for a Brownian particle we must know the Langevin equation governing the evolution of the particle. For a particle moving in the presence of a potnetial V(x) the Langevin equations are

$$\frac{\mathrm{d}x}{\mathrm{d}t} = v$$

$$\frac{\mathrm{d}v}{\mathrm{d}t} = -\frac{\gamma}{m}v + \frac{1}{m}F(x) + \frac{1}{m}\xi(t)$$
(7.2)

where the force F(x) = -V'(x). Inserting (7.2) into (7.1) gives

$$\frac{\partial}{\partial t}\varrho(x,v,t) = -\frac{\partial}{\partial x}\left(v\varrho(x,v,t)\right) + \frac{\gamma}{m}\frac{\partial}{\partial v}\left(v\varrho(x,v,t)\right) - \frac{1}{m}F(x)\frac{\partial}{\partial v}\varrho(x,v,t) \\ - \frac{1}{m}\xi(t)\frac{\partial}{\partial v}\varrho(x,v,t) = -L_0\varrho(x,v,t) - L_1(t)\varrho(x,v,t)$$

where the differential operators L_0 and L_1 are defined as

$$L_{0} = v \frac{\partial}{\partial x} - \frac{\gamma}{m} - \frac{\gamma}{m} v \frac{\partial}{\partial v} + \frac{1}{m} F(x) \frac{\partial}{\partial v}$$
$$L_{1} = \frac{1}{m} \xi(t) \frac{\partial}{\partial v}$$

Since $\xi(t)$ is a stochastic variable the time evolution of ρ will be different for each realization of $\xi(t)$. However when we observe an actual Brownian particle we are observing the average effect of the random force on it. Therefore we introduce an observable probability density $P(x, v, t) = \langle \xi(t) \rangle_{\xi}$.

Let

$$\varrho(t) = \mathrm{e}^{-L_0 t} \sigma(t)$$

then

$$\frac{\partial}{\partial t}\sigma(x,v,t) = -e^{-L_0 t}L_1(t)e^{-L_0 t}\sigma(x,v,t) = -V(t)\sigma(x,v,t)$$

7.2 Probability flow for Brownian particle

This equation has the formal solution

$$\sigma(t) = \exp\left[-\int_0^t \mathrm{d}t' V(t')\right]\sigma(0)$$

which follows since formally

$$\begin{aligned} \sigma(t) &= \sigma(0) - \int_0^t \mathrm{d}t_1 V(t_1) \sigma(t_1) = \sigma(0) - \int_0^t \mathrm{d}t_1 V(t_1) \sigma(0) + \int_0^t \mathrm{d}t_1 \int_0^{t_1} \mathrm{d}t_2 V(t_1) V(t_2) \sigma(0) + \dots \\ &+ (-1)^n \int_0^t \mathrm{d}t_1 \int_0^{t_1} \mathrm{d}t_2 \cdots \int_0^{t_n} \mathrm{d}t_{n-1} V(t_1) V(t_2) \dots V(t_{n-1}) \sigma(0) + \dots \\ &= \sum_{n=0}^\infty \frac{(-1)^n}{n!} \left[\int_0^t \mathrm{d}t_1 V(t_1) \right]^n \sigma(0) = \exp\left[-\int_0^t \mathrm{d}t' V(t') \right] \sigma(0) \end{aligned}$$

The third step follows since by changing the order of integration and then varibles

$$\int_0^t \mathrm{d}t_1 \int_0^{t_1} \mathrm{d}t_2 V(t_1) V(t_2) = \int_0^t \mathrm{d}t_2 \int_{t_2}^t \mathrm{d}t_1 V(t_1) V(t_2) = \int_0^t \mathrm{d}t_1 \int_{t_1}^t \mathrm{d}t_2 V(t_1) V(t_2)$$

so that

$$\int_0^t \mathrm{d}t_1 \int_0^{t_1} \mathrm{d}t_2 V(t_1) V(t_2) = \frac{1}{2} \left[\int_0^t \mathrm{d}t_1 V(t_1) \right]^2$$

Also assume that

$$\int_{0}^{t} \mathrm{d}t_{1} \int_{0}^{t_{1}} \mathrm{d}t_{2} \cdots \int_{0}^{t_{n}} \mathrm{d}t_{n-1} V(t_{1}) V(t_{2}) \dots V(t_{n-1}) = \frac{1}{n!} \left[\int_{0}^{t} \mathrm{d}t_{1} V(t_{1}) \right]^{n}$$
(7.3)

then by taking the derivative

$$\frac{\mathrm{d}}{\mathrm{d}t} \int_0^t \mathrm{d}t_1 \int_0^{t_1} \mathrm{d}t_2 \cdots \int_0^{t_{n+1}} \mathrm{d}t_n V(t_1) V(t_2) \dots V(t_n)$$

$$= V(t) \int_0^t \mathrm{d}t_2 \int_0^{t_2} \mathrm{d}t_3 \cdots \int_0^{t_n} \mathrm{d}t_{n+1} V(t_1) V(t_2) \dots V(t_n)$$

$$= V(t) \frac{1}{n!} \left[\int_0^t \mathrm{d}t_1 V(t_1) \right]^n = \frac{\mathrm{d}}{\mathrm{d}t} \frac{1}{(n+1)!} \left[\int_0^t \mathrm{d}t_1 V(t_1) \right]^{n+1}$$

By induction it therefore follows that (7.3) holds. Taking the average $\langle \cdots \rangle_{\xi}$ over the Gaussian noice $\xi(t)$ we see that $\langle \sigma(t) \rangle_{\xi}$ is the characteristic function of the random variable $X(t) = i \int_0^t dt_1 V(t_1)$. This must again be a Gaussian variable with $\langle X(t) \rangle_{\xi} = 0$ and the variance is

$$\langle X(t)^2 \rangle = \frac{1}{2} \int_0^t \mathrm{d}t_1 \int_0^t \mathrm{d}t_2 \langle V(t_1)V(t_2) \rangle$$

Since the characteristic function for the Gaussian variable X(t) is $\exp(iX(t)) = \exp(i\mu_X - \langle X(t)^2 \rangle/2)$ we find

$$\langle \sigma(t) \rangle_{\xi} = \exp\left(\frac{1}{2} \int_0^t \mathrm{d}t_1 \int_0^t \mathrm{d}t_2 \langle V(t_1)V(t_2) \rangle\right) \sigma(0) \tag{7.4}$$

This formula is just a special case of a cumulant expansion. The integral in (7.4) becomes

$$\frac{1}{2} \int_{0}^{t} \mathrm{d}t_{1} \int_{0}^{t} \mathrm{d}t_{2} \langle V(t_{1})V(t_{2})\rangle_{\xi} = \frac{1}{2} \int_{0}^{t} \mathrm{d}t_{1} \int_{0}^{t} \mathrm{d}t_{2} \langle \mathrm{e}^{L_{0}t_{1}} \frac{1}{m} \xi(t_{1}) \frac{\partial}{\partial v} \mathrm{e}^{-L_{0}t_{1}} \mathrm{e}^{L_{0}t_{2}} \frac{1}{m} \xi(t_{2}) \frac{\partial}{\partial v} \mathrm{e}^{-L_{0}t_{2}} \rangle_{\xi} \\
= \frac{g}{2m^{2}} \int_{0}^{t} \mathrm{d}t_{1} \mathrm{e}^{L_{0}t_{1}} \frac{\partial^{2}}{\partial v^{2}} \mathrm{e}^{-L_{0}t_{1}}$$

Then

$$\frac{\partial}{\partial t} \langle \sigma(x, v, t) \rangle_{\xi} = \frac{g}{2m^2} e^{L_0 t} \frac{\partial^2}{\partial v^2} e^{-L_0 t} \langle \sigma(x, v, t) \rangle_{\xi}$$

This gives for $\langle \varrho(x, v, t) \rangle_{\xi}$

$$\frac{\partial}{\partial t} \langle \varrho(x,v,t) \rangle_{\xi} = -L_0 \langle \varrho(x,v,t) \rangle_{\xi} + \frac{g}{2m^2} \frac{\partial^2}{\partial v^2} \langle \varrho(x,v,t) \rangle_{\xi}$$

and for the probability distribution

$$\frac{\partial}{\partial t}P(x,v,t) = -v\frac{\partial}{\partial x}P(x,v,t) - \frac{\partial}{\partial v}\left[\left(\frac{\gamma}{m}v - \frac{1}{m}F(x)\right)P(x,v,t)\right] + \frac{g}{2m^2}\frac{\partial^2}{\partial v^2}P(x,v,t)$$
(7.5)

This is the Fokker-Planck equation for the probability Pdxdv to find the Brownian particle in the interval (x, x + dx, (v, v + dv)) at time t.

We can write the Fokker-Planck equation as a continuity equation

$$\frac{\partial}{\partial t}P(x,v,t) = -\nabla \cdot \boldsymbol{j}$$

where $\nabla = \mathbf{e}_x \partial / \partial x + \mathbf{e}_v \partial / \partial v$ and the probability current is

$$\boldsymbol{j} = \boldsymbol{e}_x v P - \boldsymbol{e}_v \left[\left(\frac{\gamma}{m} v - \frac{1}{m} F(x) \right) P + \frac{g}{2m^2} \frac{\partial}{\partial v} P \right]$$

7.3 General Fokker-Planck equation

We can obtain the Fokker-Planck equation for a quite general Langevin equation for the dynamics of a set of fluctuating variables

$$\boldsymbol{a} = \{a_1, a_2, \ldots\}$$

We assume a general friction term $\nu_j(a_1, a_2, ...) = \nu_j(a)$ and assume a Gaussian noise $\xi_j(t)$ where

$$egin{array}{rcl} \langle \xi_j(t)
angle &=& 0 \ \langle \xi_i(t_2)\xi_j(t_1)
angle &=& g_{ij}\delta(t_2-t_1) \end{array}$$

The Langevin equation becomes

$$\frac{\mathrm{d}}{\mathrm{d}t}a_j(t) = \nu_j(a) + \xi_j(t)$$

or in vector form

$$\frac{\mathrm{d}}{\mathrm{d}t}\boldsymbol{a}(t) = \boldsymbol{\nu}(\boldsymbol{a}) + \boldsymbol{\xi}(t)$$

We ask for the probability distribution

$$P(\boldsymbol{a},t) = \langle \varrho(\boldsymbol{a},t) \rangle_{\xi}$$

Again from conservation of probability

$$\frac{\partial}{\partial t}\varrho(\boldsymbol{a},t) + \frac{\partial}{\partial \boldsymbol{a}} \cdot \left[\frac{\mathrm{d}\boldsymbol{a}}{\mathrm{d}t}\varrho(\boldsymbol{a},t)\right] = 0$$

Usin the Langevin equation to solve for da/dt we find

$$\frac{\partial}{\partial t}\varrho(\boldsymbol{a},t) = \frac{\partial}{\partial \boldsymbol{a}} \cdot (\nu(\boldsymbol{a})\varrho(\boldsymbol{a},t)) - \frac{\partial}{\partial \boldsymbol{a}} \cdot (\xi(t)\varrho(\boldsymbol{a},t)) = -[L_0 + L_1(t)] \varrho(\boldsymbol{a},t)$$

where

$$L_0 = \left(\frac{\partial}{\partial a} \cdot \nu(a)\right) + \nu(a) \cdot \frac{\partial}{\partial a}$$
$$L_1(t) = \xi(t) \cdot \frac{\partial}{\partial a}$$

Following the steps as above we find

$$\frac{\partial}{\partial t}P(\boldsymbol{a},t) = -\frac{\partial}{\partial \boldsymbol{a}} \cdot (\nu(\boldsymbol{a})P(\boldsymbol{a},t)) + \frac{1}{2}\frac{\partial}{\partial \boldsymbol{a}} \cdot \boldsymbol{g} \cdot \frac{\partial}{\partial \boldsymbol{a}}P(\boldsymbol{a},t)$$
(7.6)

Here \boldsymbol{g} is a tensor with elements g_{ij} .

Example

Our previous result can be obtained as a special case of (7.6). The Langevin equations are

$$\frac{\mathrm{d}x}{\mathrm{d}t} = v$$

$$\frac{\mathrm{d}v}{\mathrm{d}t} = -\frac{\gamma}{m}v + \frac{1}{m}F(x) + \frac{1}{m}\xi(t)$$

where

$$\langle \xi(t_2)\xi(t_1)\rangle = 2\gamma k_{\rm B}T\delta(t_2 - t_1)$$

Then

$$\boldsymbol{a} = \begin{pmatrix} x \\ y \end{pmatrix}, \qquad \nu(\boldsymbol{a}) = \begin{pmatrix} v \\ -\frac{\gamma}{m}v + \frac{1}{m}F(x) \end{pmatrix}$$
$$\boldsymbol{\xi}(t) = \begin{pmatrix} 0 \\ \frac{1}{m}\boldsymbol{\xi}(t) \end{pmatrix}, \qquad \boldsymbol{g} = \begin{pmatrix} 0 & 0 \\ 0 & \frac{2\gamma k_{\mathrm{B}}T}{m^{2}} \end{pmatrix}$$

The Fokker-Planck equation becomes

$$\frac{\partial}{\partial t}P(x,v,t) = -\frac{\partial}{\partial x}\left[vP(x,v,t)\right] - \frac{\partial}{\partial v}\left[\left(-\frac{\gamma}{m}v + \frac{1}{m}F(x)\right)P(x,v,t)\right] + \frac{\gamma k_{\rm B}T}{m^2}\frac{\partial^2}{\partial v^2}P(x,v,t)$$